

Cosmology

Part A (problems to be handed in)

1. Consider a Universe with pressureless matter, a cosmological constant, and spatial curvature. Show that the 1st Friedmann equation can be written as the equation of motion of a particle moving in one dimension with total energy zero and potential $V(a)$ as follows (where $a=R/R_0$, with $R_0=1$)

$$V(a) = -\frac{4\pi G}{3} \frac{\rho_{m,0}}{a} + \frac{k}{2} - \frac{\Lambda}{6} a^2 \quad (7 \text{ points})$$

2. Show that the age of the Universe in a spatially flat model $\Omega_{\Lambda,0} + \Omega_{m,0} = 1$ is given by

$$t(z) = \frac{2}{3H_0\sqrt{\Omega_{\Lambda,0}}} \ln \left(\sqrt{\frac{\Omega_{\Lambda,0}}{\Omega_{m,0}} \left(\frac{1}{1+z}\right)^3} + \sqrt{\frac{\Omega_{\Lambda,0}}{\Omega_{m,0}} \left(\frac{1}{1+z}\right)^3 + 1} \right) \quad (6 \text{ points})$$

3. Show that $(x = \Omega_m(a), y = \Omega_\Lambda(a))$ satisfy the homogeneous system of equations:

$$x' = -x(1 - x + 2y)$$

$$y' = y(x + 2 - 2y)$$

where $(\prime) = d/d \log a$. Compute the critical points and draw the cosmic flow/phase space diagram in the plane (x, y) .

(6 points)

4. Show that our Universe has *not* suffered a bounce since decoupling because the maximum redshift z_{\max} would satisfy $z_{\max}^2(z_{\max} + 3) < 2/\Omega_{m,0}$ but we have certainly observed galaxies at $z > 8$.

hint: A bounce corresponds to a Universe that has suffered first contraction and then expansion.

(6 points)

Part B (problems to be discussed in class)

1) Derive the expression $a=1/(1+z)$

2) Show that $\rho R^{3(1+w)} = \text{const.}$ for barotropic fluids under adiabatic expansion

3) Derive the evolution of the density parameter for a fluid with $w(z)$

4) Derive the evolution of the scale factor $a(t)$ with time during matter and radiation domination, i.e. $a(t) \sim t^n$, $n=1/2$ or $2/3$

5) Same as before, but with conformal time η defined as $d\eta = dt/a(t)$

general remarks:

- ‘show’ means that you should derive the equations rather than verifying their correctness; this also means that you have to give every single step of calculation and not just say ‘as we can see’.
- if you refer to some calculations or results from somewhere else you must provide the proper reference.
- be clear about what you do and write
- for all the exercise you should ignore any contribution coming from radiation, i.e. set $\Omega_r=0$.