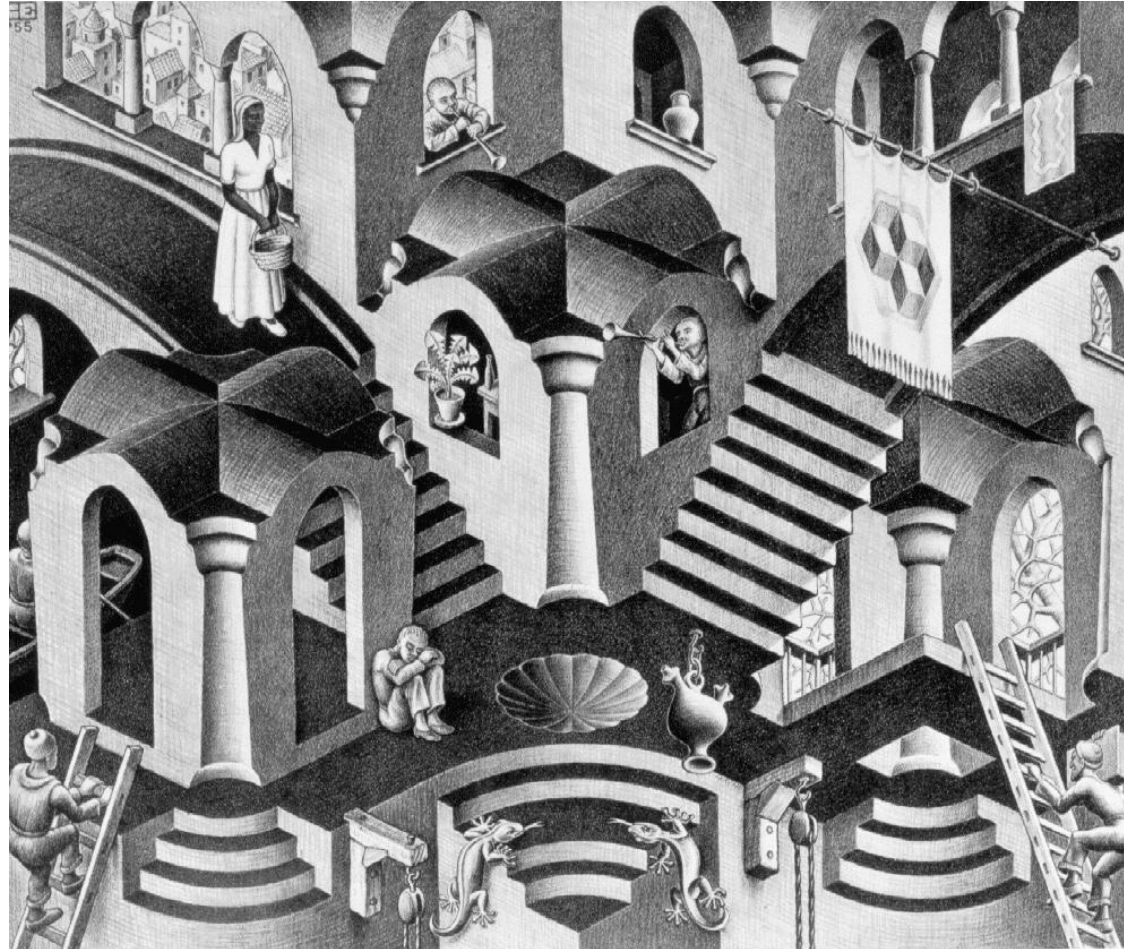


Computational Astrophysics

Solving for Gravity

Alexander Knebe, *Universidad Autonoma de Madrid*



Solving for Gravity

- Poisson's equation

$$\Delta\Phi(\vec{r}) = 4\pi G\rho(\vec{r})$$

$$\vec{F}(\vec{r}) = -m\nabla\Phi(\vec{r})$$

Solving for Gravity

- Poisson's equation

$$\Delta\Phi(\vec{r}) = 4\pi G\rho(\vec{r})$$

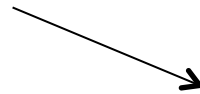
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Solving for Gravity

- Poisson's equation

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grid approach ($\vec{r}_{i,j,k}$ = position of centre of grid cell (i,j,k))

$$\Delta\Phi(\vec{r}_{i,j,k}) = 4\pi G\rho(\vec{r}_{i,j,k})$$

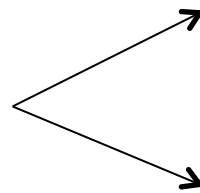
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Solving for Gravity

- Poisson's equation

$$\Delta\Phi(\vec{r}) = 4\pi G\rho(\vec{r})$$

$$\vec{F}(\vec{r}) = -m\nabla\Phi(\vec{r})$$



particle approach

$$\vec{F}(\vec{r}_i) = -\sum_{i \neq j} \frac{Gm_i m_j}{(r_i - r_j)^3} (\vec{r}_i - \vec{r}_j)$$

grid approach ($\vec{r}_{i,j,k}$ = position of centre of grid cell (i,j,k))

$$\Delta\Phi(\vec{r}_{i,j,k}) = 4\pi G\rho(\vec{r}_{i,j,k})$$

$$\vec{F}(\vec{r}_{i,j,k}) = -m\nabla\Phi(\vec{r}_{i,j,k})$$

Solving for Gravity

- direct particle-particle summation (PP)

$$\vec{F}_i(\vec{r}_i) = - \sum_{i \neq j} \frac{Gm_i m_j}{(r_i - r_j)^3} (\vec{r}_i - \vec{r}_j) \quad \forall i \in N$$

organizing particles into a “tree structure” will give $N \log(N)$ operations

Solving for Gravity

- direct particle-particle summation (PP)

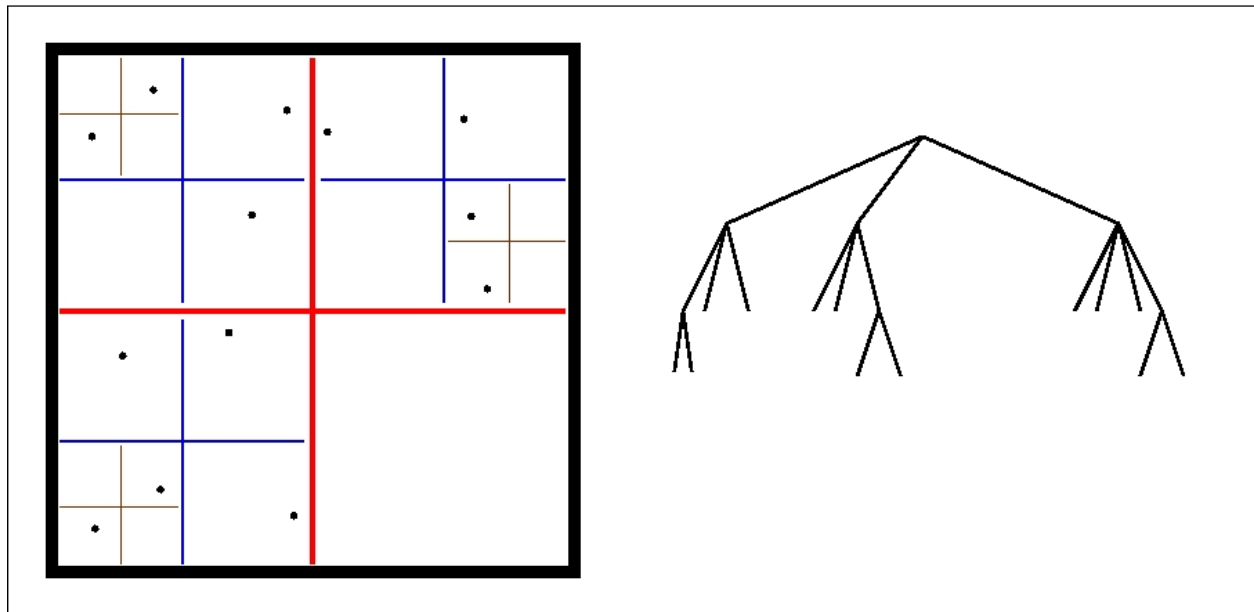
the tree

Solving for Gravity

- direct particle-particle summation (PP)

$$\vec{F}_i(\vec{r}_i) = - \sum_{i \neq j} \frac{Gm_i m_j}{(r_i - r_j)^3} (\vec{r}_i - \vec{r}_j) \quad \forall i \in N$$

- generating the tree:

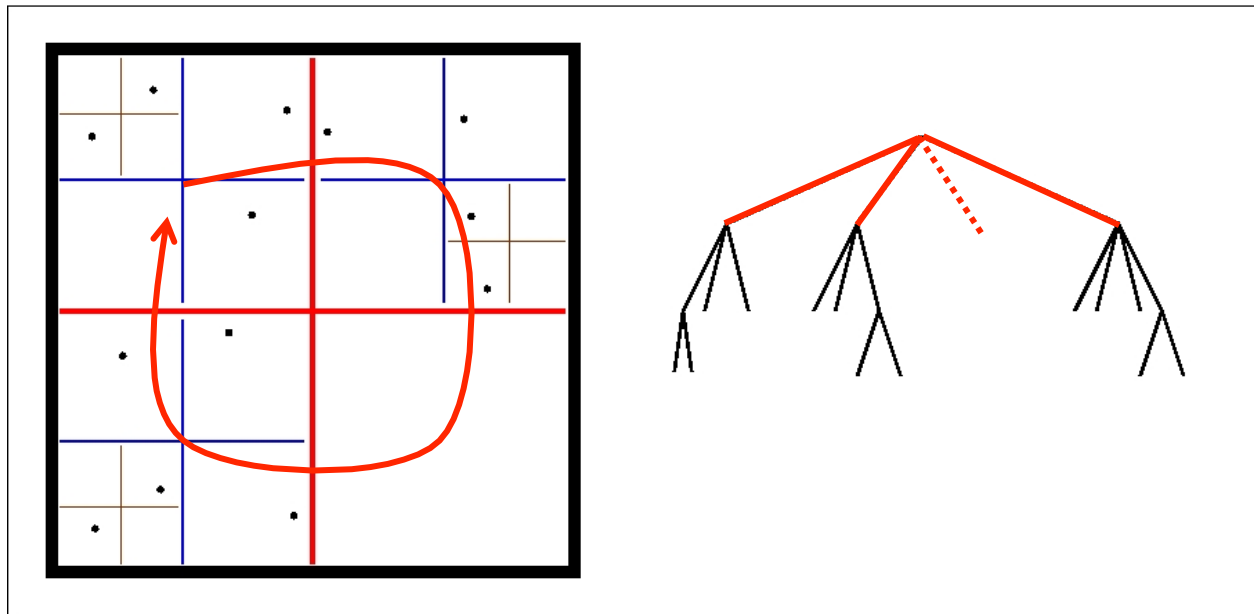


Solving for Gravity

- direct particle-particle summation (PP)

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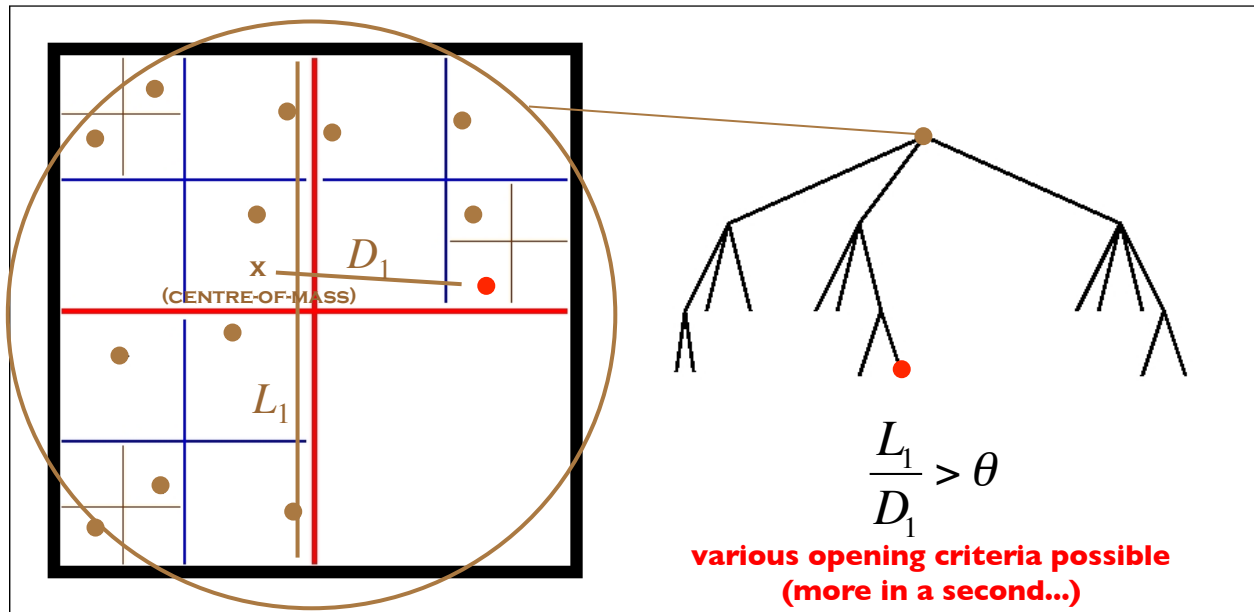


Solving for Gravity

- direct particle-particle summation (PP)

$$\vec{F}_i(\vec{r}_i) = - \sum_{i \neq j} \frac{Gm_i m_j}{(r_i - r_j)^3} (\vec{r}_i - \vec{r}_j) \quad \forall i \in N$$

- walking the tree ($\forall i \in N$):



Solving for Gravity

requirements to perform
Cosmological Simulations

Solving for Gravity

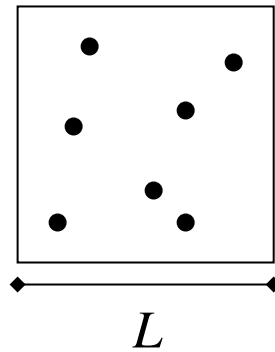
- specific requirements for cosmological simulations:
 - periodic boundary conditions
 - equations-of-motion in comoving coordinates

Solving for Gravity

- specific requirements for cosmological simulations:
 - **periodic boundary conditions**
 - equations-of-motion in comoving coordinates

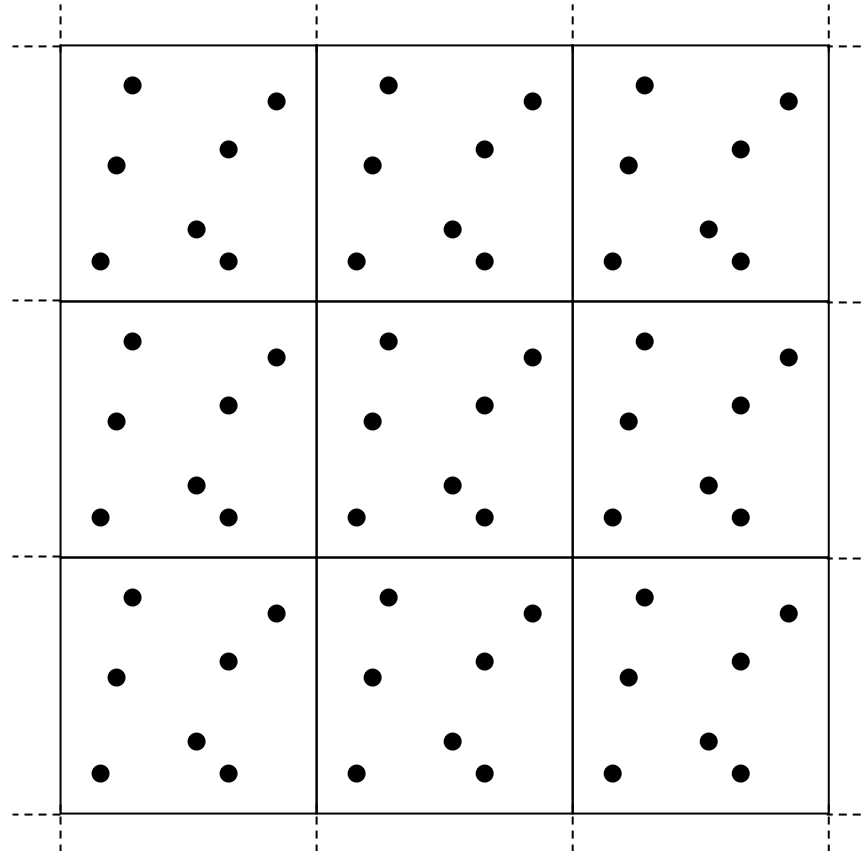
Solving for Gravity

- periodic boundary conditions



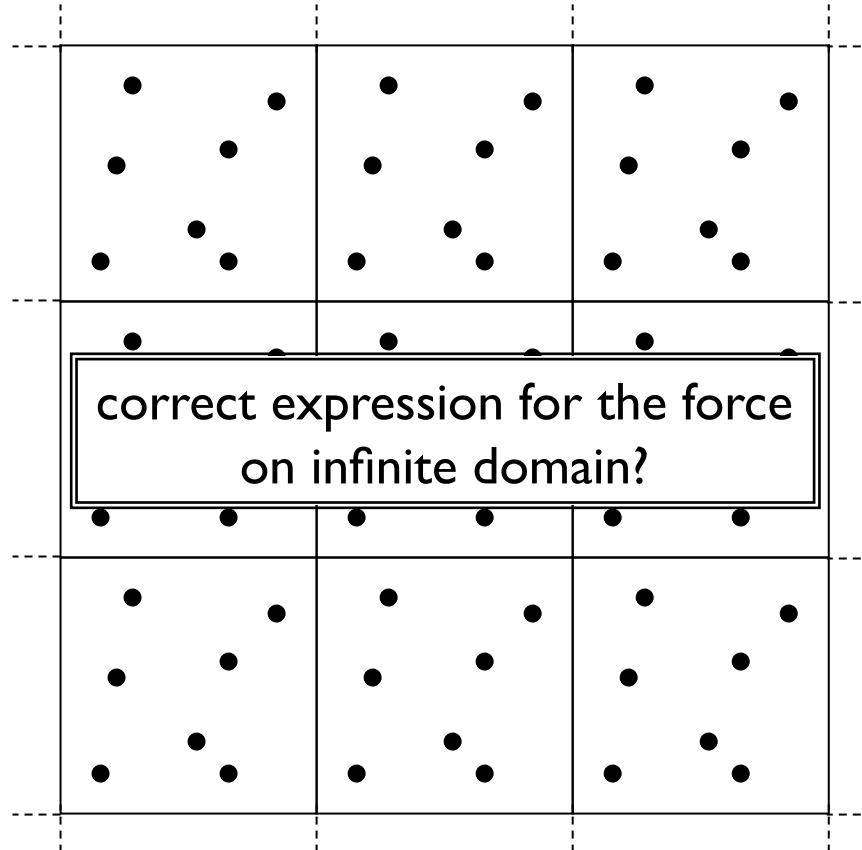
Solving for Gravity

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Solving for Gravity

- periodic boundary conditions



Solving for Gravity

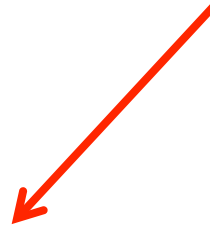
- periodic boundary conditions

$$\Delta_x \Phi(\vec{x}) = 4\pi G(\rho_x(\vec{x}) - \bar{\rho}_x)$$

Solving for Gravity

- periodic boundary conditions

$$\Delta_x \Phi(\vec{x}) = 4\pi G(\rho_x(\vec{x}) - \bar{\rho}_x)$$



we need to subtract the mean background density
in order for the solution to converge!

Solving for Gravity

- periodic boundary conditions

$$\Delta_x \Phi(\vec{x}) = 4\pi G(\rho_x(\vec{x}) - \bar{\rho}_x)$$



general solution

$$\Phi(\vec{x}) = G \iiint \frac{\rho_x(\vec{x}') - \bar{\rho}_x}{|\vec{x} - \vec{x}'|} d^3 x'$$

Solving for Gravity

- periodic boundary conditions

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periodicity automatically taken care of when using PM solver!

Solving for Gravity

- periodic boundary conditions

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periodicity automatically taken care of when using PM solver!

...but needs to be properly implemented for tree codes!

Solving for Gravity

- periodic boundary conditions

$$\Delta_x \Phi(\vec{x}) = 4\pi G(\rho_x(\vec{x}) - \bar{\rho}_x)$$



general solution

fluctuates about zero!

$$\Phi(\vec{x}) = G \iiint \frac{\rho_x(\vec{x}') - \bar{\rho}_x}{|\vec{x} - \vec{x}'|} d^3 x'$$

$$i.e., \Phi(\vec{0}) = -G \iiint \frac{(\rho_x(\vec{x}) - \bar{\rho}_x)}{|\vec{x}|} d^3 x$$

$$= -G \int_{x=0}^{\infty} \frac{1}{|\vec{x}|} x^2 \iint_{\vartheta, \varphi} (\rho_x(\vec{x}) - \bar{\rho}_x) \sin \vartheta d\vartheta d\varphi$$

$$= -G \int_{x=0}^{\infty} x \langle \rho_x(\vec{x}) - \bar{\rho}_x \rangle_{|\vec{x}|=x} dx$$

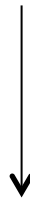
⇒ convergence to **finite value**, as $\langle \rangle \rightarrow 0$ for $x \rightarrow \infty$

Note: what is true to $x=0$ is true for any point as the origin of the coordinate system is arbitrary for periodic boundaries...

Solving for Gravity

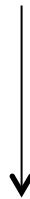
- periodic boundary conditions

$$\Delta_x \Phi(\vec{x}) = 4\pi G(\rho_x(\vec{x}) - \bar{\rho}_x)$$



general solution

$$\Phi(\vec{x}) = G \iiint \frac{\rho_x(\vec{x}') - \bar{\rho}_x}{|\vec{x} - \vec{x}'|} d^3 x'$$



desired (peculiar) force field

$$\vec{F}(\vec{x}) = -G \iiint \frac{\rho_x(\vec{x}') - \bar{\rho}_x}{|\vec{x} - \vec{x}'|^3} (\vec{x} - \vec{x}') d^3 x'$$

Solving for Gravity

- periodic boundary conditions

$$\vec{F}(\vec{x}) = -G \iiint \frac{\rho_x(\vec{x}') - \bar{\rho}_x}{|\vec{x} - \vec{x}'|^3} (\vec{x} - \vec{x}') d^3 x'$$

Solving for Gravity

- periodic boundary conditions

...but in the end it will not contribute to F !

$$\vec{F}(\vec{x}) = -G \iiint \frac{\rho_x(\vec{x}') - \bar{\rho}_x}{|\vec{x} - \vec{x}'|^3} (\vec{x} - \vec{x}') d^3 x'$$

Solving for Gravity

- periodic boundary conditions

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$$\begin{aligned} i.e., \vec{F}(\vec{0}) &= -G \iiint \frac{(\rho_x(\vec{x}) - \bar{\rho}_x)}{|\vec{x}|^3} \vec{x} d^3 x \\ &= -G \iiint \frac{\rho_x(\vec{x})}{|\vec{x}|^3} \vec{x} d^3 x + G \iiint \frac{\bar{\rho}_x}{|\vec{x}|^3} \vec{x} d^3 x \end{aligned}$$

$$\iiint \frac{\vec{x}}{|\vec{x}|^3} d^3 x = \iiint_{x, \vartheta, \varphi} \frac{1}{x^3} \begin{pmatrix} x \cos \varphi \sin \vartheta \\ x \sin \varphi \sin \vartheta \\ x \cos \vartheta \end{pmatrix} x^2 \sin \vartheta dx d\vartheta d\varphi = \iiint_{x, \vartheta, \varphi} \begin{pmatrix} \cos \varphi \sin \vartheta \\ \sin \varphi \sin \vartheta \\ \cos \vartheta \end{pmatrix} \sin \vartheta dx d\vartheta d\varphi = \begin{pmatrix} 0 \\ 0 \\ 0 \end{pmatrix}$$

Solving for Gravity

- periodic boundary conditions

$$\vec{F}(\vec{x}) = -G \iiint \frac{\rho_x(\vec{x}')}{|\vec{x} - \vec{x}'|^3} (\vec{x} - \vec{x}') d^3 x'$$

Solving for Gravity

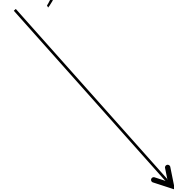
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particle/discrete picture

$$\vec{F}(\vec{x}) = -G \sum_{i=1}^N \sum_{\vec{R}} \frac{m_i}{|\vec{x} - (\vec{x}_i + \vec{R})|^3} (\vec{x} - (\vec{x}_i + \vec{R}))$$



$$\vec{R} = \vec{n}L$$

Solving for Gravity

- periodic boundary conditions

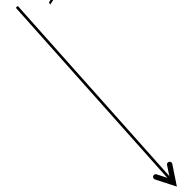
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particle/discrete picture

$$\vec{F}(\vec{x}) = -G \sum_{i=1}^N \sum_{\vec{R}} \frac{m_i}{|\vec{x} - (\vec{x}_i + \vec{R})|^3} (\vec{x} - (\vec{x}_i + \vec{R}))$$

correct expression for the force on infinite domain!

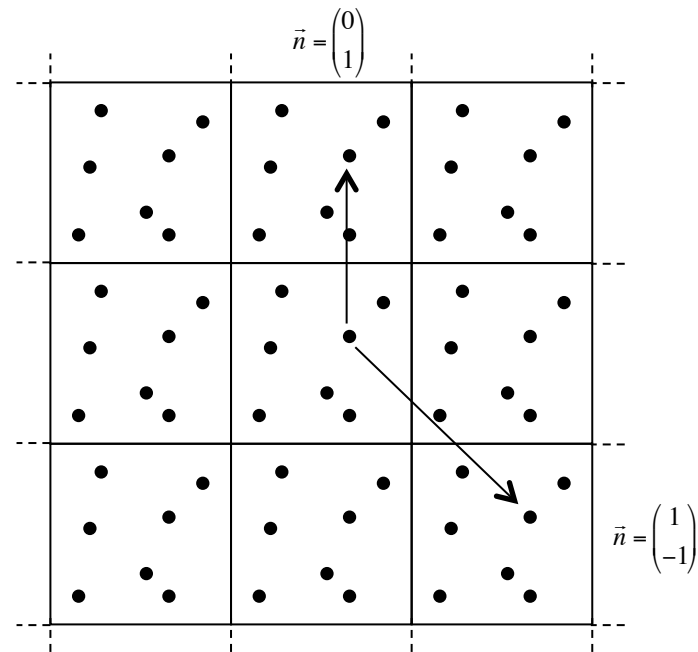


$$\vec{R} = \vec{n}L$$

Solving for Gravity

- periodic boundary conditions

$$\vec{F}(\vec{x}) = -G \sum_{i=1}^N \sum_{\vec{R}} \frac{m_i}{|\vec{x} - (\vec{x}_i + \vec{R})|^3} (\vec{x} - (\vec{x}_i + \vec{R}))$$



Solving for Gravity

- periodic boundary conditions

$$\vec{F}(\vec{x}) = -G \sum_{i=1}^N \sum_{\vec{R}} \frac{m_i}{|\vec{x} - (\vec{x}_i + \vec{R})|^3} (\vec{x} - (\vec{x}_i + \vec{R}))$$

=> slow convergence and hence not feasible...

=> Ewald summation instead...

Solving for Gravity

- periodic boundary conditions

$$\Delta_x \Phi(\vec{x}) = 4\pi G \rho(\vec{x})$$

Solving for Gravity

- periodic boundary conditions

$$\Delta_x \Phi(\vec{x}) = 4\pi G \rho(\vec{x})$$

$$\rho(\vec{x}) = \sum_{i=1}^N m_i \delta_{\text{Dirac}}(\vec{x} - \vec{x}_i)$$



“peculiar” density

$$\rho(\vec{x}) = \sum_{i=1}^N m_i \delta_{\text{Dirac}}(\vec{x} - \vec{x}_i) - \bar{\rho}$$



periodic, peculiar density

$$\rho_{\text{periodic}}(\vec{x}) = \sum_{i=1}^N \sum_{\vec{R}} m_i \delta_{\text{Dirac}}(\vec{x} - (\vec{x}_i + \vec{R})) - \bar{\rho}$$

$$\vec{R} = \vec{n}L \quad (\vec{n} = \text{integer vector})$$

Solving for Gravity

- periodic boundary conditions

$$\rho_{\text{periodic}}(\vec{x}) = \sum_{i=1}^N \sum_{\vec{R}} m_i \delta_{\text{Dirac}}(\vec{x} - \vec{x}_i - \vec{R}) - \bar{\rho}$$

Solving for Gravity

- periodic boundary conditions

$$\rho_{\text{periodic}}(\vec{x}) = \sum_{i=1}^N \sum_{\vec{R}} m_i \delta_{\text{Dirac}}(\vec{x} - \vec{x}_i - \vec{R}) - \bar{\rho}$$


$$\rho_{\text{periodic}}(\vec{x}) = \sum_{i=1}^N \rho_1(\vec{x}, \vec{x}_i) + \rho_2(\vec{x}, \vec{x}_i)$$

Solving for Gravity

- periodic boundary conditions

$$\rho_{\text{periodic}}(\vec{x}) = \sum_{i=1}^N \rho_1(\vec{x}, \vec{x}_i) + \rho_2(\vec{x}, \vec{x}_i)$$

$$\rho_1(\vec{x}, \vec{x}_i) = \sum_{\vec{R}} m_i \delta_{\text{Dirac}}(\vec{x} - \vec{x}_i - \vec{R})$$

$$\rho_2(\vec{x}, \vec{x}_i) = -\bar{\rho}$$

Solving for Gravity

- periodic boundary conditions

$$\rho_{\text{periodic}}(\vec{x}) = \sum_{i=1}^N \rho_1(\vec{x}, \vec{x}_i) + \rho_2(\vec{x}, \vec{x}_i)$$

$$\rho_1(\vec{x}, \vec{x}_i) = - \sum_{\vec{R}} \frac{1}{\sqrt{\mu^2 \pi}} e^{-\frac{(\vec{x} - \vec{x}_i - \vec{R})^2}{\mu^2}} + \sum_{\vec{R}} m_i \delta_{\text{Dirac}}(\vec{x} - \vec{x}_i - \vec{R})$$

$$\rho_2(\vec{x}, \vec{x}_i) = + \sum_{\vec{R}} \frac{1}{\sqrt{\mu^2 \pi}} e^{-\frac{(\vec{x} - \vec{x}_i - \vec{R})^2}{\mu^2}} - \bar{\rho}$$

Ewald introduced (Gaussian) “screening charges”

Solving for Gravity

- periodic boundary conditions

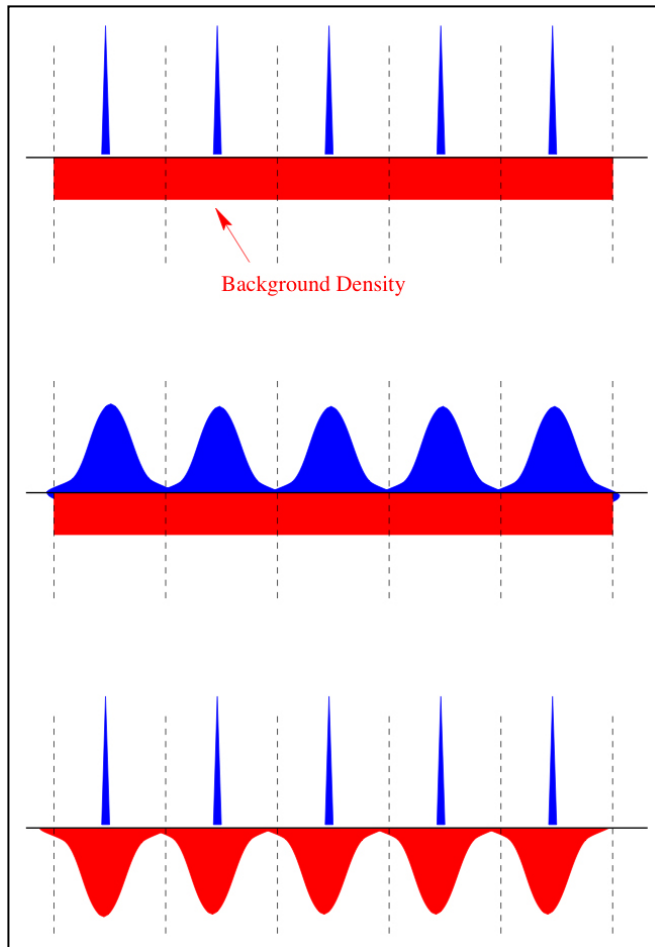
$$\rho_{\text{periodic}}(\vec{x}) = \sum_{i=1}^N \rho_1(\vec{x}, \vec{x}_i) + \rho_2(\vec{x}, \vec{x}_i)$$

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$$\rho_2(\vec{x}, \vec{x}_i) = + \sum_{\vec{R}} \frac{1}{\sqrt{\mu^2 \pi}} e^{-\frac{(\vec{x} - \vec{x}_i - \vec{R})^2}{\mu^2}} - \bar{\rho} \quad \rightarrow \text{Fourier-space}$$

Solving for Gravity

- periodic boundary conditions



$$\rho_{\text{periodic}}(\vec{x}) = \sum_{i=1}^N \rho_1(\vec{x}, \vec{x}_i) + \rho_2(\vec{x}, \vec{x}_i)$$

potential obtained in...

$$\rho_2(\vec{x}, \vec{x}_i) = \sum_{\vec{R}} \frac{1}{\sqrt{\mu^2 \pi}} e^{-\frac{(\vec{x} - \vec{x}_i - \vec{R})^2}{\mu^2}} - \bar{\rho}$$

Fourier-space

$$\rho_1(\vec{x}, \vec{x}_i) = \sum_{\vec{R}} m_i \delta_{\text{Dirac}}(\vec{x} - \vec{x}_i - \vec{R}) - \frac{1}{\sqrt{\mu^2 \pi}} e^{-\frac{(\vec{x} - \vec{x}_i - \vec{R})^2}{\mu^2}}$$

real-space

=> exponential convergence and hence feasible!

Solving for Gravity

- periodic boundary conditions

detailed calculation...

Solving for Gravity

- periodic boundary conditions

- force due to particles in computational box:

$$\vec{F}(\vec{x}) = -G \sum_{i=1}^N \frac{m_i}{|\vec{x} - \vec{x}_i|^3} (\vec{x} - \vec{x}_i)$$

- *additional* force due to periodic images:

$$\vec{F}_{Ewald}(\vec{x}) = \frac{\vec{x}}{x^3} - \sum_{\vec{R}} \frac{\vec{x} - \vec{R}}{|\vec{x} - \vec{R}|^3} \times \left[\operatorname{erfc}\left(\frac{|\vec{x} - \vec{R}|}{\mu}\right) + \frac{2|\vec{x} - \vec{R}|}{\sqrt{\mu^2 \pi}} e^{-\frac{|\vec{x} - \vec{R}|^2}{\mu^2}} \right] - \frac{2}{L^2} \sum_{\vec{n} \neq 0} \frac{\vec{n}}{n} \sin\left(\frac{2\pi}{L} \vec{n} \cdot (\vec{x} - \vec{R})\right) e^{-\frac{(\mu\pi)^2 n^2}{L^2}}$$

Ewald summation in practice...

Solving for Gravity

- periodic boundary conditions

- additional force due to periodic images:

$$\vec{F}_{Ewald}(\vec{x}) = \frac{\vec{x}}{x^3} - \sum_{\vec{R}} \frac{\vec{x} - \vec{R}}{|\vec{x} - \vec{R}|^3} \times \left[\operatorname{erfc}\left(\frac{|\vec{x} - \vec{R}|}{\mu}\right) + \frac{2|\vec{x} - \vec{R}|}{\sqrt{\mu^2 \pi}} e^{-\frac{|\vec{x} - \vec{R}|^2}{\mu^2}} \right] - \frac{2}{L^2} \sum_{\vec{n} \neq 0} \frac{\vec{n}}{n} \sin\left(\frac{2\pi}{L} \vec{n} \cdot (\vec{x} - \vec{R})\right) e^{-\frac{(\mu\pi)^2 n^2}{L^2}}$$

- in practice:

1. $\mu = L/2, \quad |\vec{x} - \vec{R}| < 3L, \quad n^2 < 10$

2. tabulate $F_{Ewald}(x)$ on a grid and interpolate...

Solving for Gravity

- specific requirements for cosmological simulations:
 - periodic boundary conditions
 - **equations-of-motion in comoving coordinates**

Solving for Gravity

(cf. "gastrophysics in supercomoving coordinates" lecture...)

$$\frac{\partial \rho_x}{\partial T} + \nabla_x \cdot (\rho_x \vec{v}) = 0$$

$$\frac{\partial(\rho_x \vec{v})}{\partial T} + \nabla_x \cdot (\rho_x \vec{v} \otimes \vec{v} + p_x \vec{1}) = \rho_x (-\nabla_x \phi_x)$$

$$\frac{\partial(\rho_x E_x)}{\partial T} + \nabla_x \cdot ([\rho_x E_x + p_x] \vec{v}) = \rho_x \vec{v} \cdot (-\nabla_x \phi_x) - \mathcal{H} \rho_x \varepsilon_x [3\gamma - 5] + (\Gamma_x - L_x)$$

$$\frac{\partial S_x}{\partial T} + \nabla_x \cdot (S_x \vec{v}) = -\mathcal{H} S_x [3\gamma - 5]$$

additional/closure equations:

$$\Delta_x \phi_x = 4\pi G a (\rho_{x,tot} - \bar{\rho}_x)$$

$$\frac{d\vec{x}_{DM}}{dT} = \vec{v}_{DM}$$

$$\frac{d\vec{v}_{DM}}{dT} = -\nabla_x \phi_x$$

$$E_x = \varepsilon_x + \frac{1}{2} v^2$$

$$p_x = (\gamma - 1) \rho_x \varepsilon_x$$

$$S_x = \frac{p_x}{\rho_x^{\gamma-1}}$$

$$\varepsilon_x = \frac{1}{(\gamma - 1)} \frac{1}{\mu m_p} T_x$$

Solving for Gravity

▪ comoving coordinates

• recap:

$$\vec{r} = a\vec{x}, \quad \vec{v} = \vec{u} + H\vec{r}$$

$$\dot{\vec{x}} = \vec{u} / a$$

$$\dot{\vec{u}} = \vec{f} - H\vec{u}$$

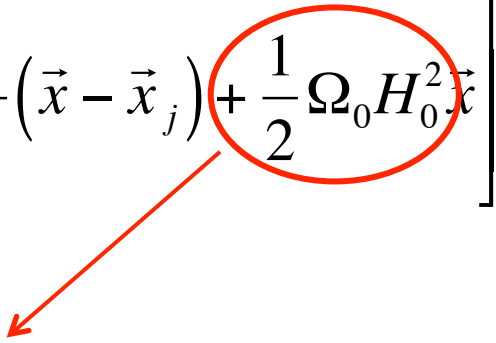
$$\vec{f} = -G \frac{1}{a^2} \sum_{x \neq x_j}^N \frac{m_j}{|\vec{x} - \vec{x}_j|^3} (\vec{x} - \vec{x}_j) - \ddot{a}\vec{x}$$

(cf. “Time Integration” lecture...)

Solving for Gravity

- comoving coordinates

- GADGET notation:

$$\vec{f} = \frac{1}{a^2} \left[-G \sum_{x \neq x_j}^N \frac{m_j}{|\vec{x} - \vec{x}_j|^3} (\vec{x} - \vec{x}_j) + \frac{1}{2} \Omega_0 H_0^2 \vec{x} \right]$$


$$\ddot{a} = -\frac{4\pi G}{3} a \rho = -\frac{4\pi G}{3} a \frac{\rho_0}{a^3} = -\frac{4\pi G}{3} \frac{\Omega_0 \rho_{crit,0}}{a^2} = -\frac{4\pi G}{3} \frac{\Omega_0}{a^2} \frac{3H_0^2}{8\pi G} = -\frac{1}{a^2} \frac{1}{2} \Omega_0 H_0^2$$

Solving for Gravity

▪ comoving coordinates

- isolated boundaries:

$$\vec{f} = -G \frac{1}{a^2} \sum_{x \neq x_j}^N \frac{m_j}{|\vec{x} - \vec{x}_j|^3} (\vec{x} - \vec{x}_j) - \ddot{a}\vec{x}$$

- periodic boundaries:

$$\vec{f} = -G \frac{1}{a^2} \sum_{x \neq x_j}^N \sum_{\vec{R}} \frac{m_j}{|\vec{x} - \vec{x}_j - \vec{R}|^3} (\vec{x} - \vec{x}_j - \vec{R})$$

(cf. discussion about periodic boundaries earlier...)